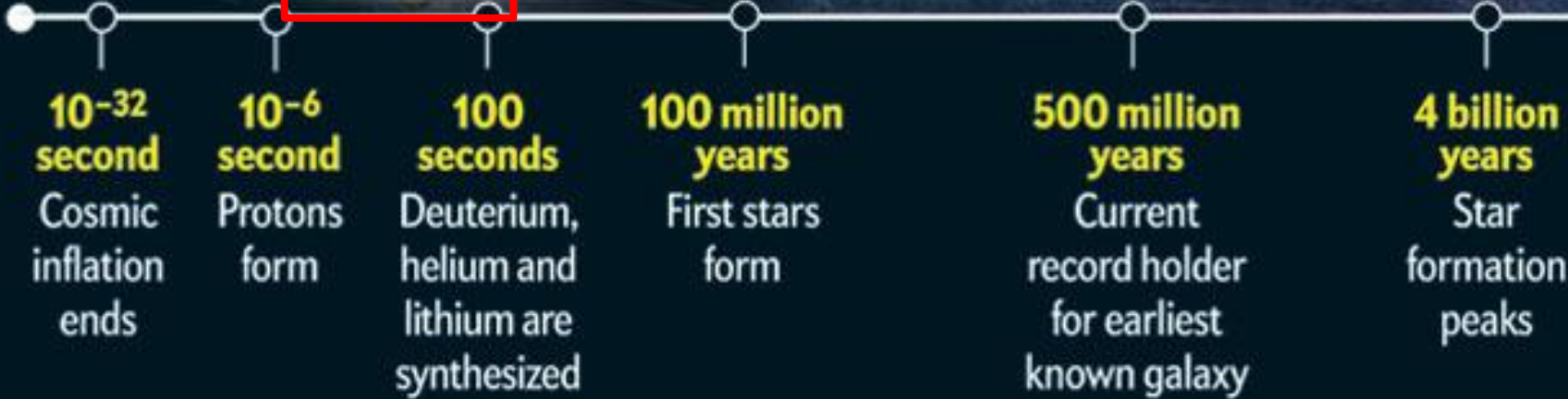
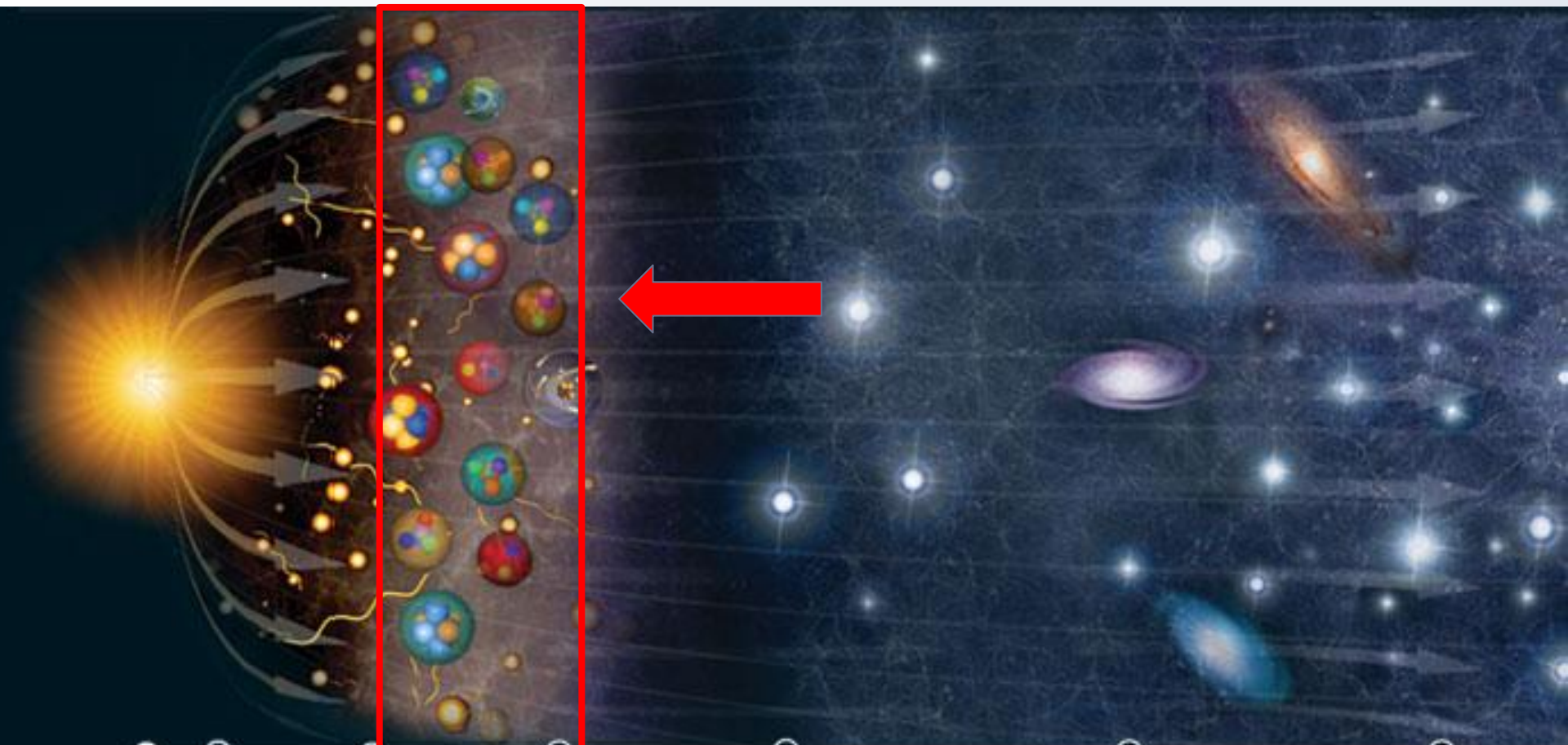


Catalyzing Interdisciplinarity

Rafael S. de Souza

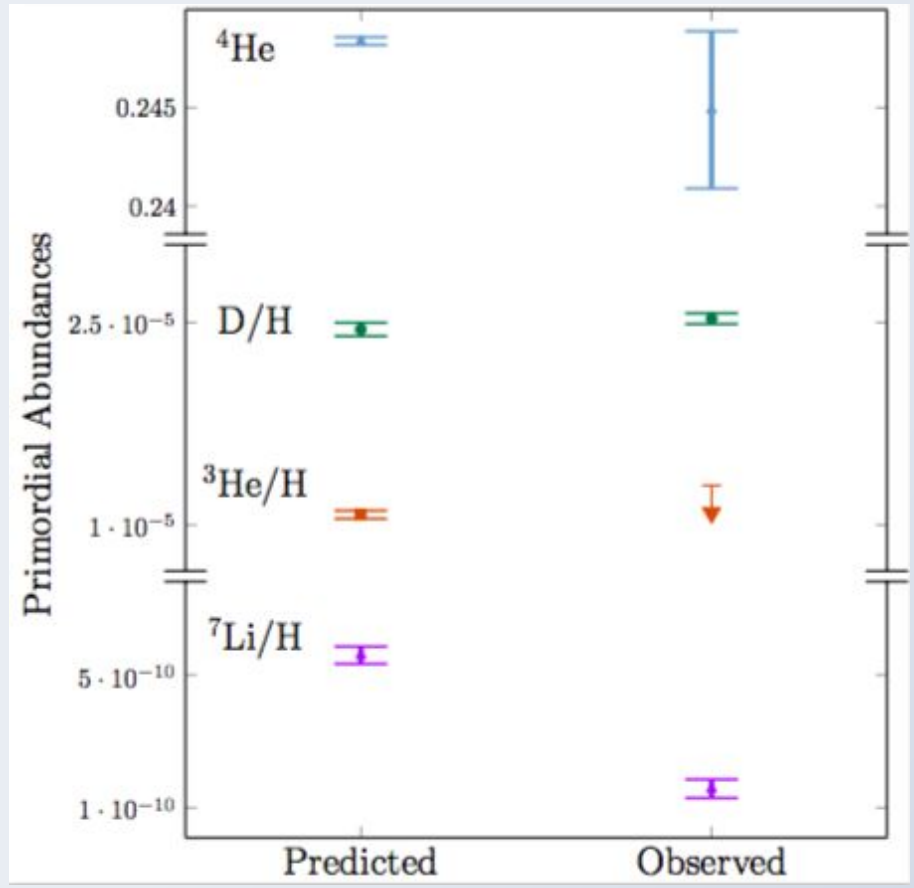
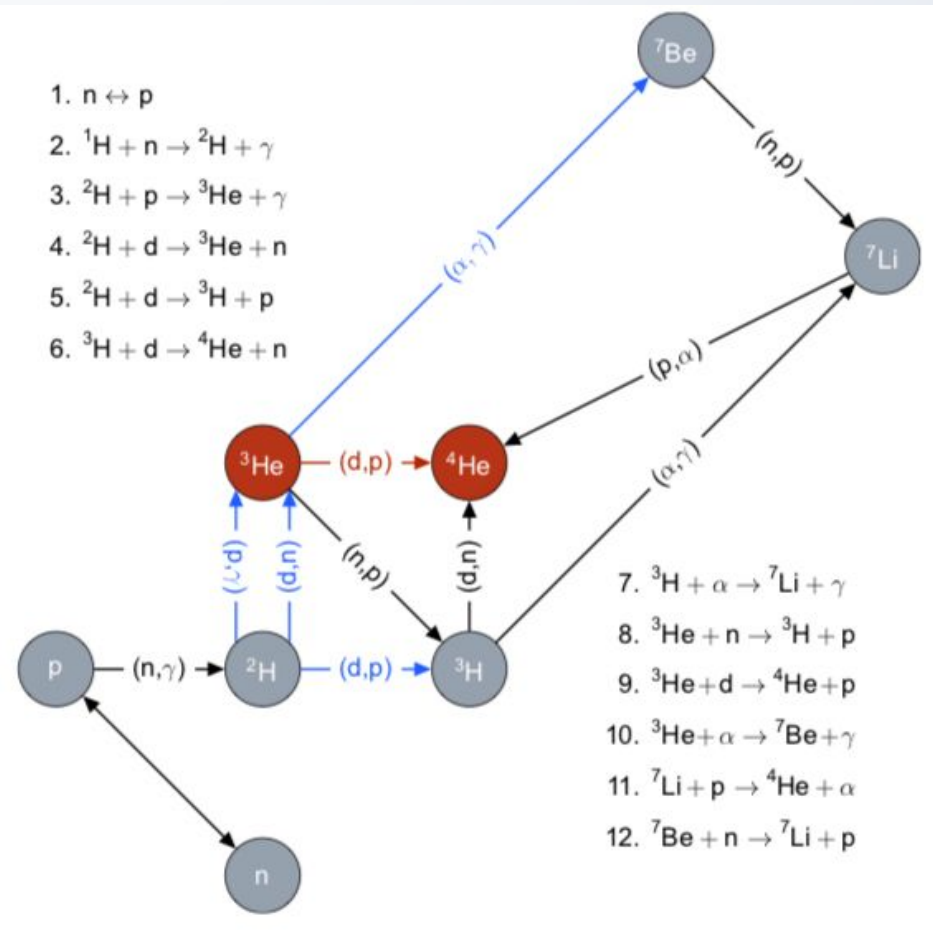
Shanghai Astronomical Observatory

Cosmostatistics Initiative



The Big-Bang Nucleosynthesis

Reaction rates  Primordial Abundances



Common sources of uncertainties in Nuclear Astrophysics

- ❑ Plasma in laboratory differs from Astrophysical Plasma
- ❑ Unknown stochastic scatter;
- ❑ Errors-in-measurements
- ❑ Systematic effects - unknown normalization factors from different experiments

Hierarchical Bayesian Thermonuclear Rate for the ${}^7\text{Be}(n,p){}^7\text{Li}$ Big Bang Nucleosynthesis Reaction

Rafael S. de Souza^{1,2} , Tan Hong Kiat³, Alain Coc⁴, and Christian Iliadis^{1,2} 

Published 2020 May 15 • © 2020. The American Astronomical Society. All rights reserved.

[The Astrophysical Journal, Volume 894, Number 2](#)

Astrophysical S -factors, Thermonuclear Rates, and Electron Screening Potential for the ${}^3\text{He}(d,p){}^4\text{He}$ Big Bang Reaction via a Hierarchical Bayesian Model

Rafael S. de Souza¹ , Christian Iliadis^{1,2} , and Alain Coc³

Published 2019 February 12 • © 2019. The American Astronomical Society. All rights reserved.

[The Astrophysical Journal, Volume 872, Number 1](#)

Uncertainty quantification for Nuclear Astrophysics via Bayesian Hierarchical Models



PHYSICAL REVIEW C

covering nuclear physics

Highlights Recent Accepted Authors Referees Search Press About 

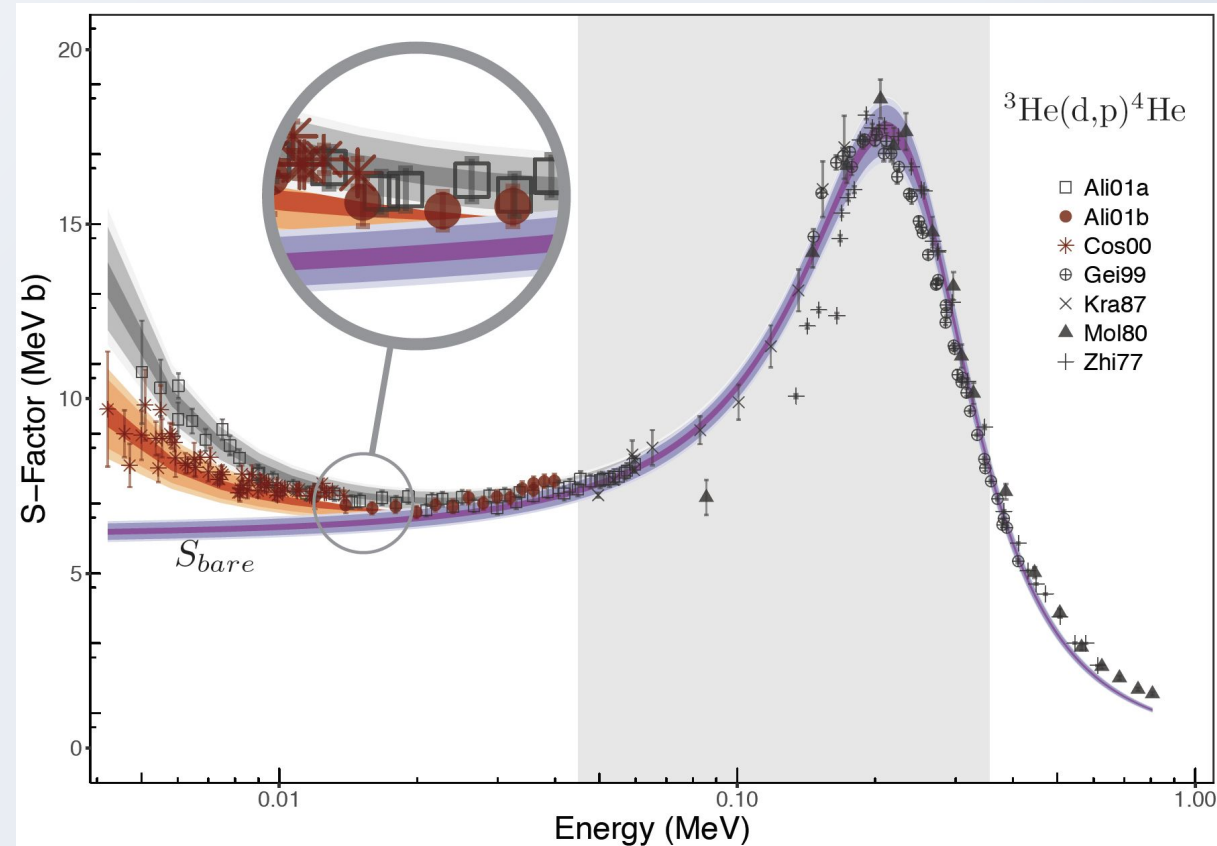
Accepted Paper

Thermonuclear fusion rates for tritium + deuterium using Bayesian methods

Phys. Rev. C

Rafael S. de Souza, S. Reece Boston, Alain Coc, and Christian Iliadis

Accepted 2 January 2019



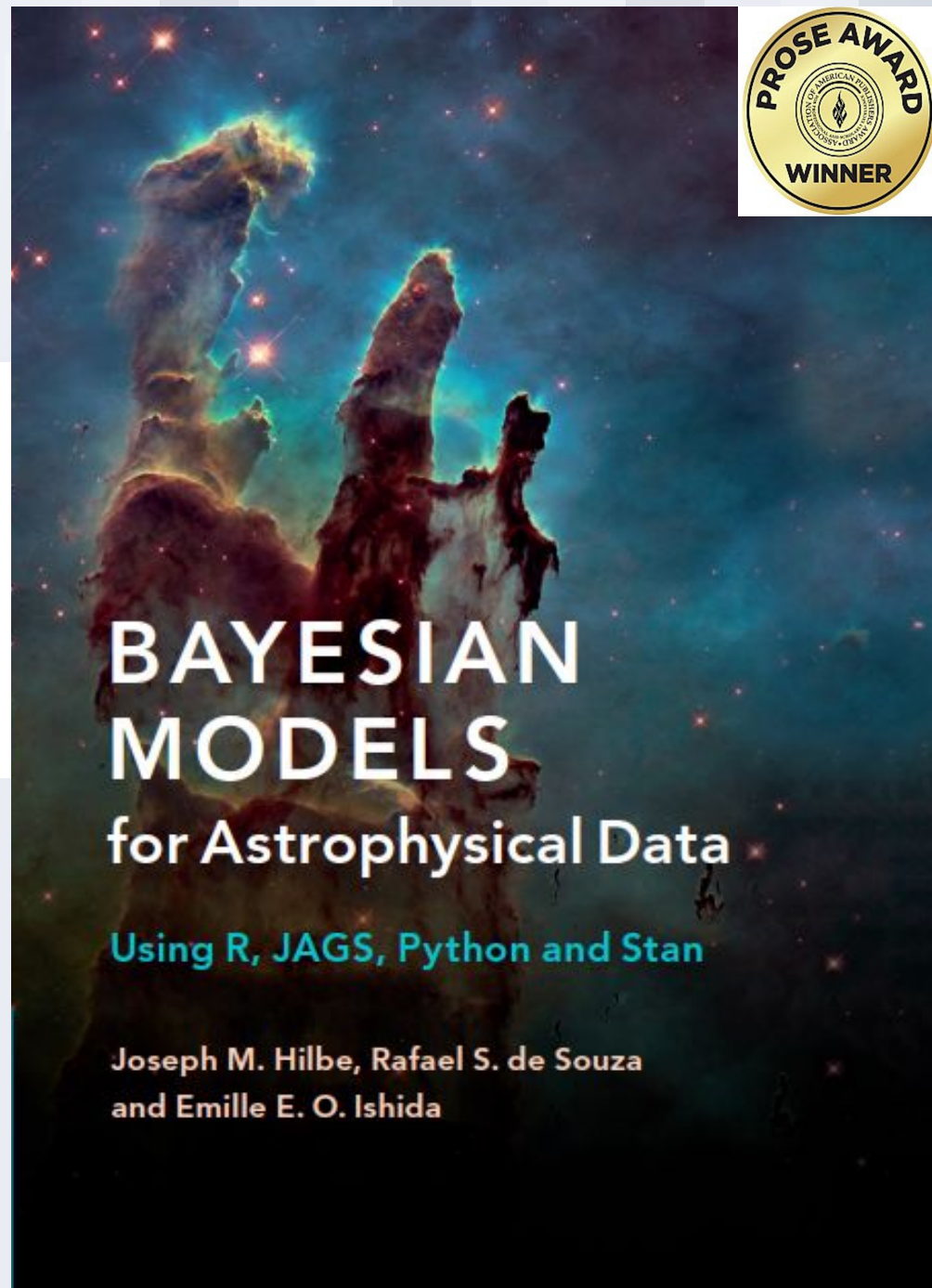
COSMOLOGY & ASTRONOMY



Cambridge University Press

*Bayesian Models for Astrophysical Data Using R, JAGS,
Python, and Stan*

By Joseph M. Hilbe, Rafael S. de Souza, and Emille E. O. Ishida



Astrostatistics

The development of new statistical methods and their application in the quest to answer astronomical endeavours has given rise, in the past few decades, to the field of Astrostatistics. This exciting area, that emerges at the intersection of Statistics and Astronomy, provides a fertile ground for both communities to grow. This innovative series from Cambridge University Press is a new medium to communicate advances in Astrostatistics to enable a steady astronomical–statistical dialogue, to develop a common language for the benefit of both communities, and to catalyse the synergy between them.

Elements are short, timely, broadly accessible papers that will appeal to both astronomical and statistical communities, highlighting cutting-edge developments for graduate students and researchers. This Elements series offers a unique platform where statistical methods and their potential applications may be demonstrated, for example, in advance of an important astronomical data release, as well as reviews and tutorials on more general topics in Astrostatistics, accepting both invited and unsolicited contributions, both subject to suitable peer review by astronomers and statisticians. The series also encourages best practice in software/code archiving and distribution through appropriate repositories to ensure their long-term access and scientific reproducibility.

Forthcoming topics in the series

(among many others planned over the coming few years):

- Sparsity in Astronomical Data Analysis and Acquisition
- Astronomical Inference via Forward Modelling and Template Libraries
- Bayesian Inference for Astrophysics
- Poisson Statistics in High-Energy Astrophysics

Series Editors

Rafael de Souza, *University of North Carolina at Chapel Hill*

Emille Ishida, *Université d'Auvergne*

Alberto Krone-Martins, *Universidade de Lisboa*

Jianhua Huang, *Texas A&M University*

Alan Heavens, *Imperial College London*

Benjamin Wandelt, *Université de Paris*

Cambridge Elements are original, concise, authoritative, and peer-reviewed collections of scholarly and scientific research. Organised into focused series edited by leading scholars, they provide comprehensive coverage of key topics in disciplines spanning the arts and sciences.

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With lengths of approx. 20,000 to 30,000 words (~40–70pp.), Elements offer an opportunity to publish concise overviews of techniques and their applications not typically found in the traditional book and journal literature.

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Elements are short; authors will be provided with a standard template, and each manuscript will be copy-edited. Elements will be published within 12 weeks of acceptance of the final manuscript.

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Relevant metadata for individual elements will be sent to the key abstracting and indexing organisations, and Altmetric data for individual Elements will also be available.

Open Access

In order to comply with the requirements of funding agencies, Cambridge will offer Open Access publication options for Elements.

Functionality

The digital-first format allows:

- Enhanced search
- Linked references
- Citation export
- Ability to link to a variety of associated formats – audio, video
- Richer content – social media, debates, resources for teaching, links to data repositories

Content will be hosted on Cambridge Core, which means Elements will benefit from launch marketing as well as increased discoverability. Functionality will be updated on a regular basis.

Want to find out more?



To write an Element for this series or to find out more information about it, contact:
Rafael de Souza (rafael@cosmostatistics-initiative.org)



For more information about the Cambridge Elements publishing model, contact Vince Higgs (vhiggs@cambridge.org)



www.cambridge.org/elements

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Cosmostatistics Initiative

A worldwide endeavour aimed to foster interdisciplinary collaborations around Astronomy.

Team Science Learning



STATISTICAL CHALLENGES in 21st CENTURY COSMOLOGY

IAU SYMPOSIUM 306 Lisbon Portugal 25-29 May 2014



Session: CMB (Chair: Graca Rocha)

16h15 – Anomalies – Hiranya Peiris

16h50 – Transforming Data into Science: Planck data and the CMB non-Gaussianity – Anna Mangilli

17h10 – Applications of the Gaussian Kinematic Formula in Cosmology – Yabebal Fantaye

17h30 – Detectability of multi-connected topologies – Ophélie Fabre

17h50 – Cosmology with photometric quasars – Boris Leistedt

18h10 – Session ends

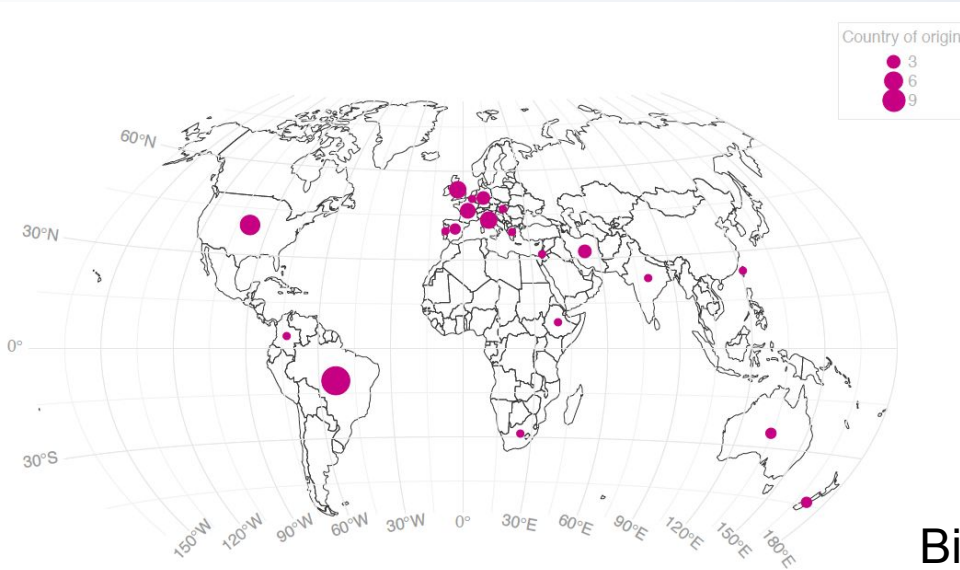
18h10 to 18h40 – *Meeting of the IAA Working Group on Cosmostatistics – Hosted by Rafael de Souza*



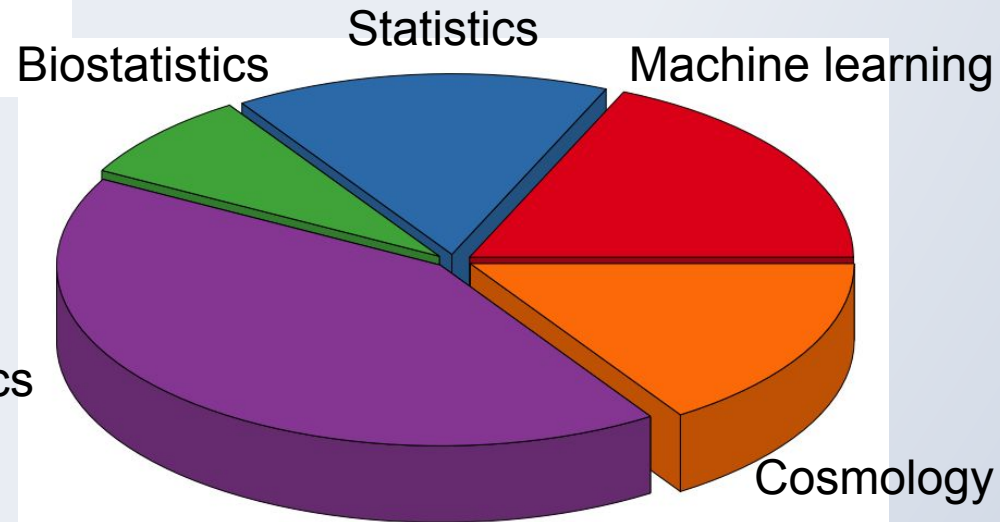
A worldwide task force

Management model utilizes concepts of startups and meta-studies of interdisciplinary teams.

Projects, and the structure of the venue, are designed to optimize the contribution of each participant



Over **60** researchers from **15** countries



See e.g. [Am J Prev Med.](#) 2008, 35, S96-115. *The ecology of team science understanding contextual influences on transdisciplinary collaboration*

C*IN*

Cosmostatistics Initiative



<https://cosmostatistics-initiative.org>

About

Chairs

Projects

Residence Programs

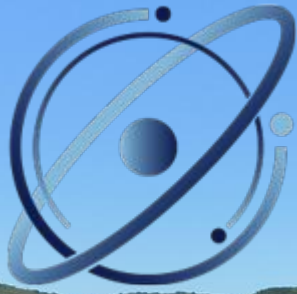
Focus

Community

Highlights

Partners

CRP#4: Clermont-Ferrand, France 2017



CRP#5: Chania, Greece 2018



COIN Residence Program #6
Chamonix - France, 24 - 31 August 2019





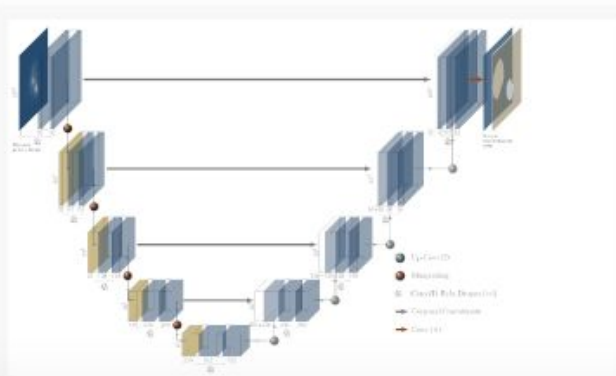
COIN-Focus : TOLIMAN mission

[Read more](#)



COIN and LSST-DESC join forces in astro-wise machine learning research

[Read more](#)



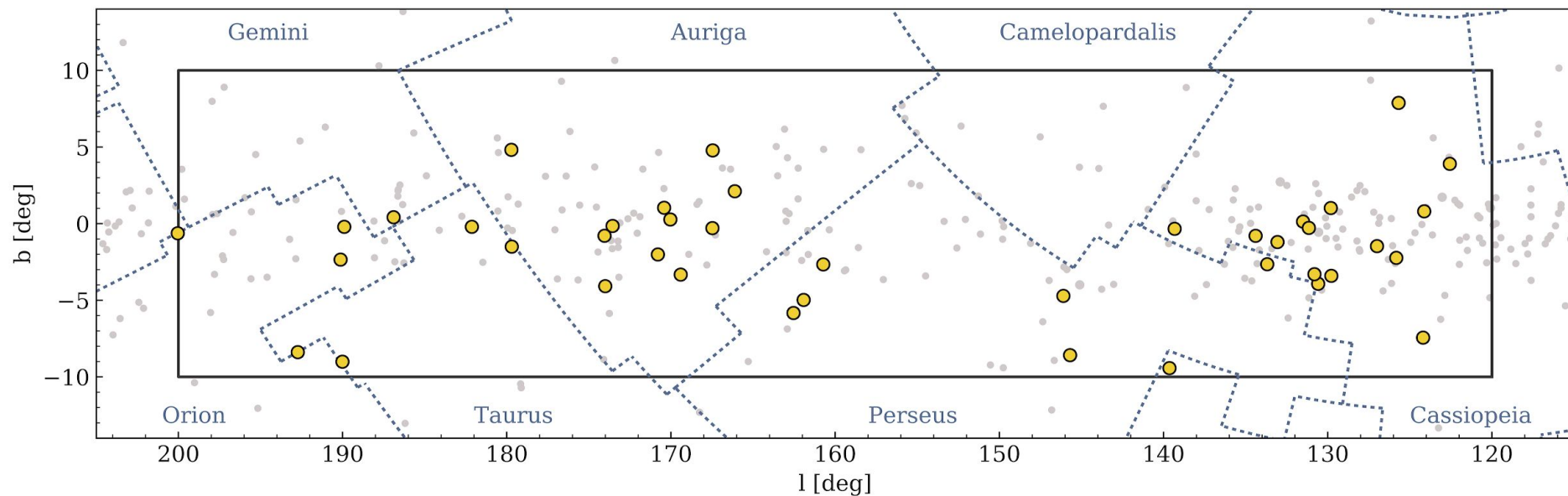
Sight beyond sight: Teasing Galaxies Apart with Deep Learning

[Read more](#)



Now you see me: COIN extends the open cluster census in the solar neighborhood with Gaia DR2

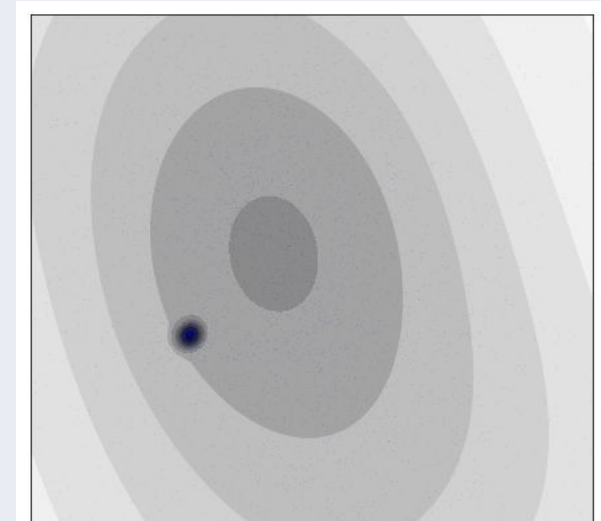
A&A, 624, A126, 17, 2019



Astro-aware statistical learning



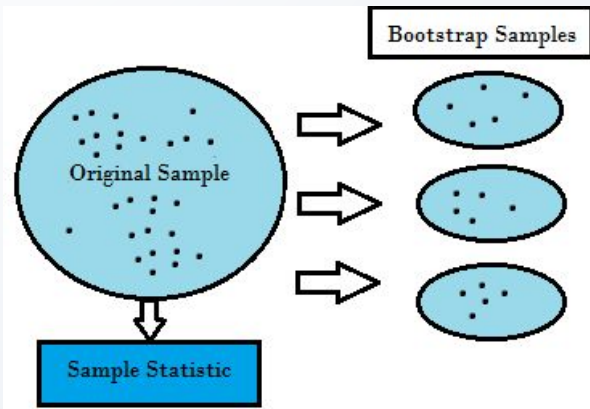
Fast density-aware partitioning of the sky via a k -d tree in the spatial domain of Galactic coordinates.



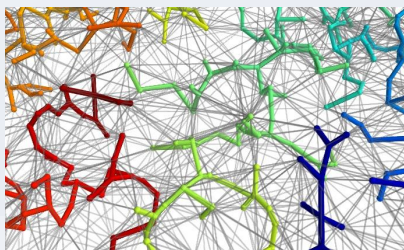
Only meaningful cases are further scrutinized, i.e. low variance in proper motion.

Recommender system

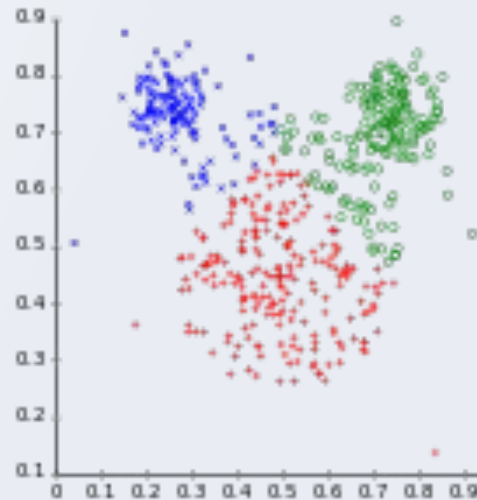
Astro-aware statistical learning



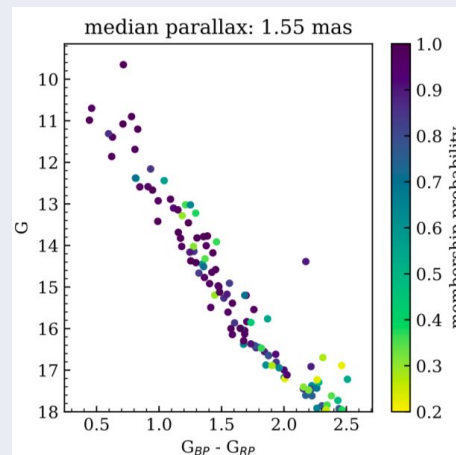
Bootstrap for measurement uncertainty



Sanity check against a random field via minimum spanning trees

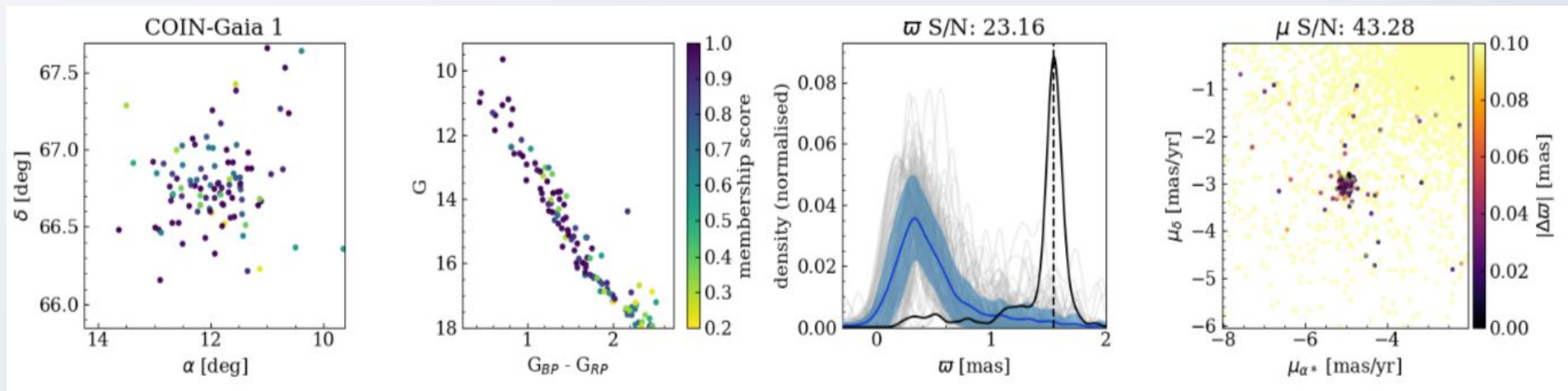


Iterative K-means in the space of proper motions.



Independent expert validation

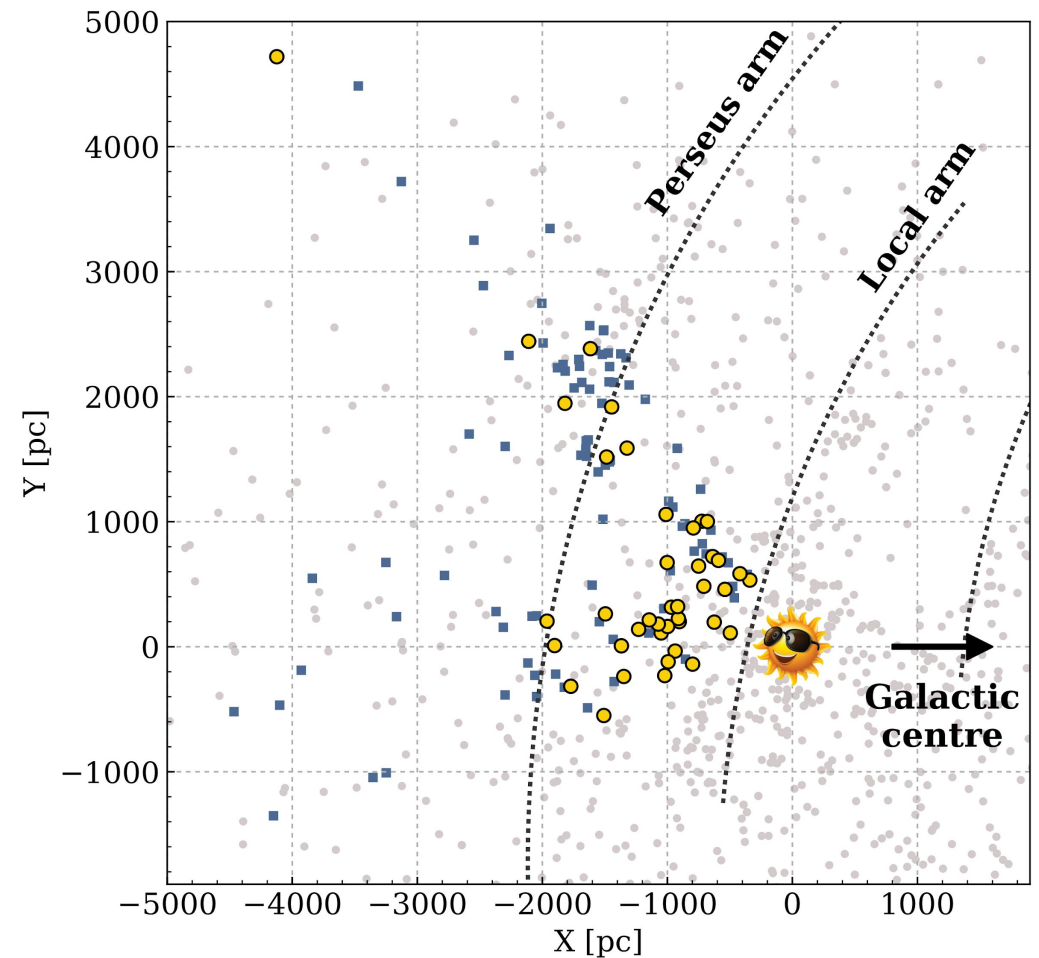
COIN-Gaia 1



COIN extends the open cluster census in the solar neighborhood with Gaia DR2



We reported the discovery of 45 new stellar clusters

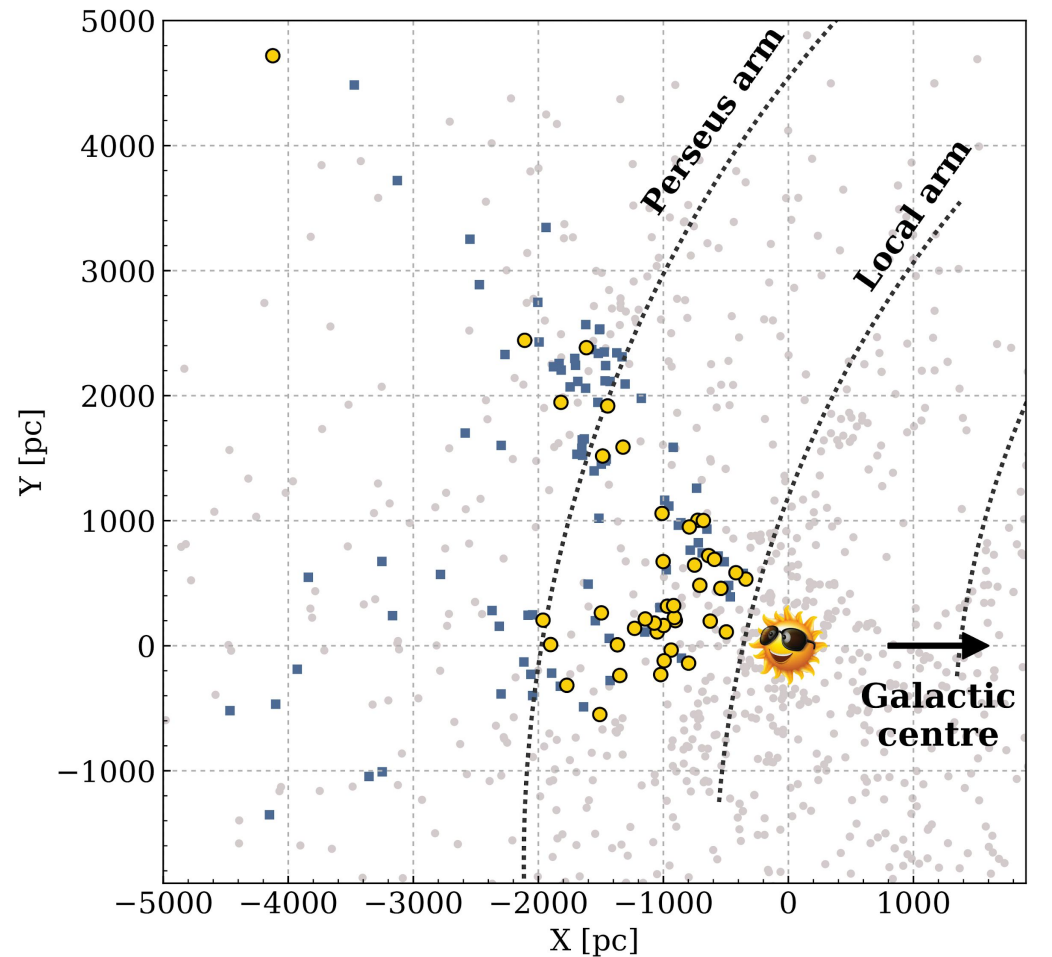


An increment of at least 20% of the previously known OC population in this volume of the Milky Way.

Potential follow-up project- GAIA DR3 + LAMOST:



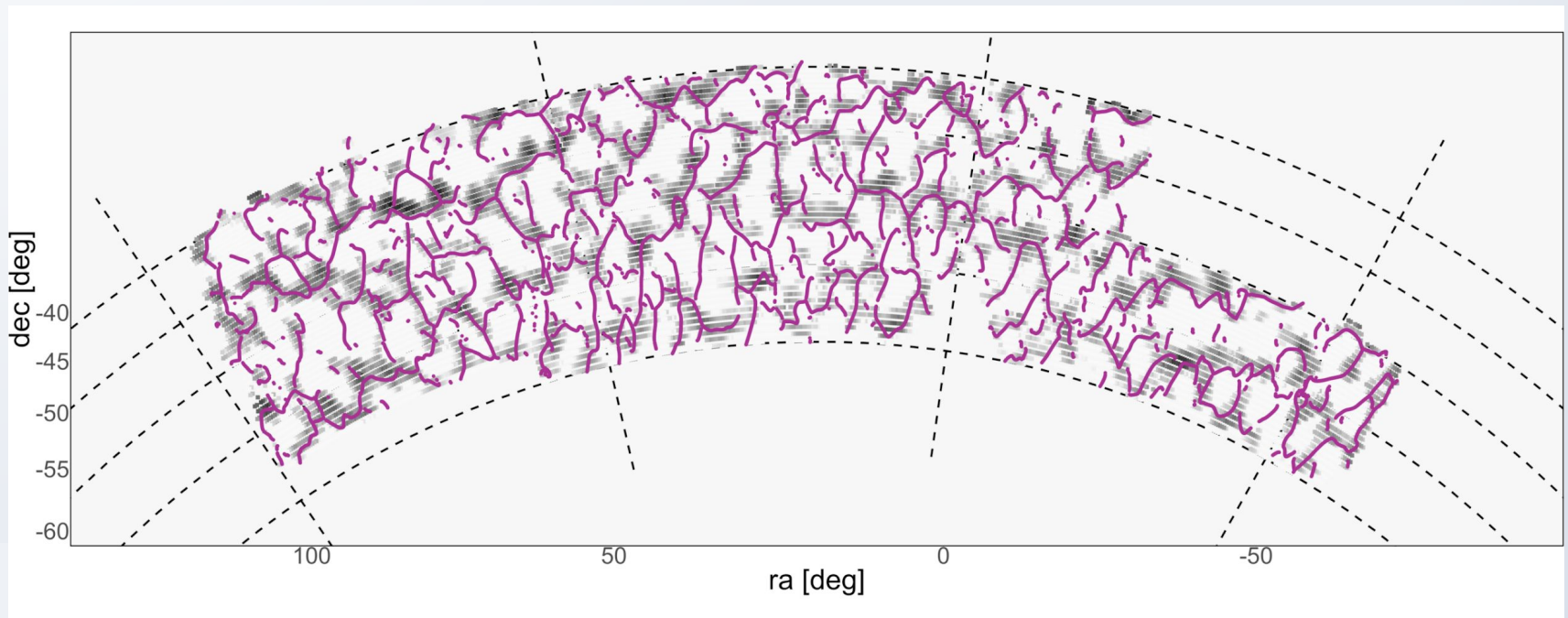
Full sky search with Gaia DR3 and cross-match the newly discovered with **LAMOST-LEGUE** and related **spectroscopic surveys** to produce the largest OC catalog of OC properties.



An increment of at least 20% of the previously known OC population in this volume of the Milky Way.

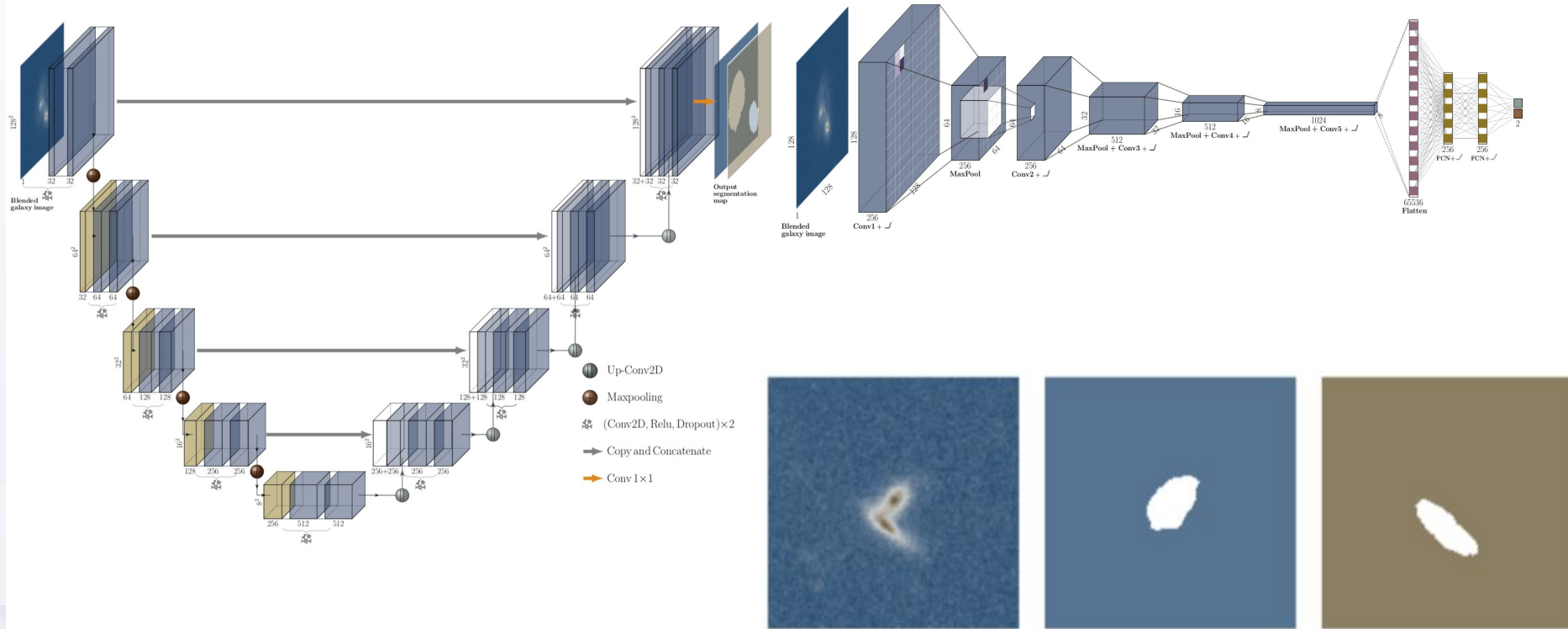
Ridges in the Dark Energy Survey for cosmic voids/trough identification

Cosmic voids/troughs play an important role in our attempt to model the large-scale structure of the Universe. They are probes to alternative cosmologies, and dark energy equation of state.



Extension of the subspace-constrained mean shift algorithm applied to 2D weak-lensing density maps from the DES DR1 [arXiv:2005.08583](https://arxiv.org/abs/2005.08583).

Deep Learning for Galaxy Deblending



From **COIN Residence Program #5**, MNRAS, 491, 2481 (2020)

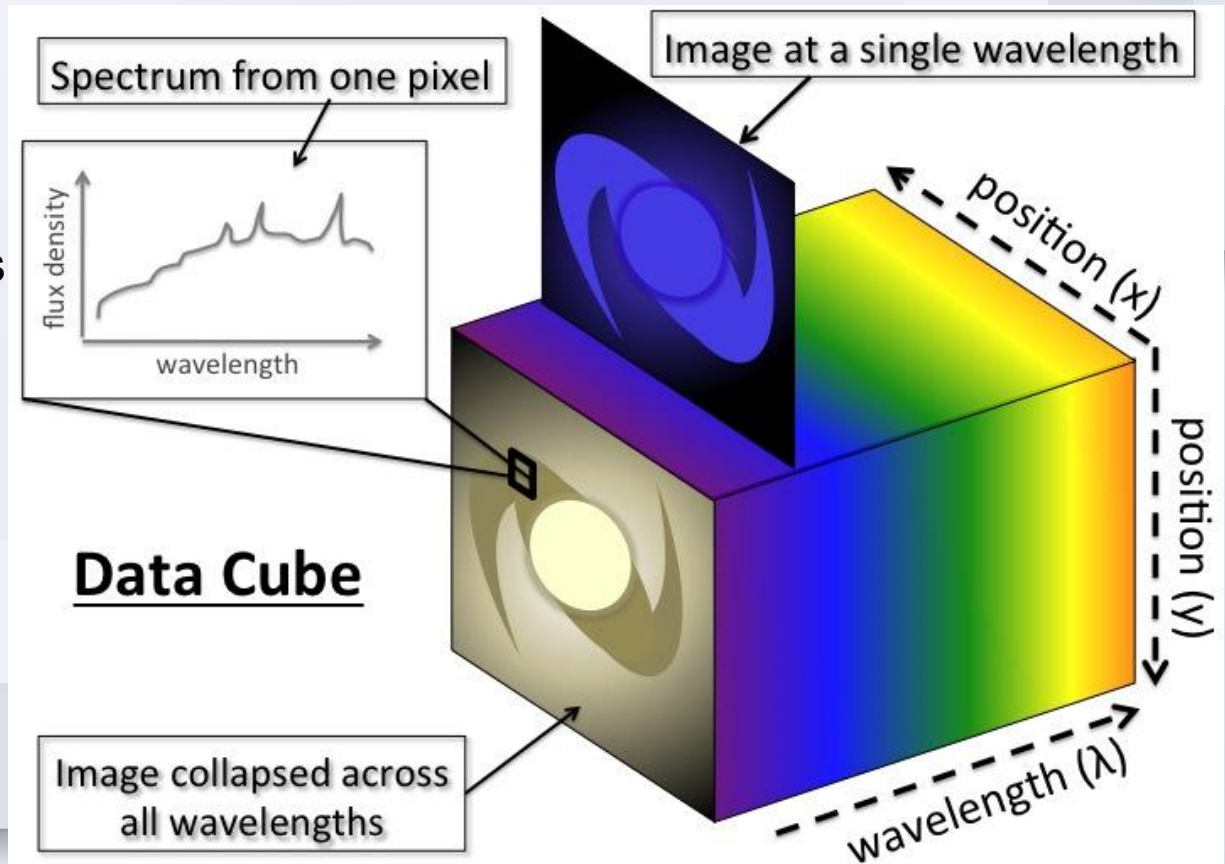


Spatial field reconstruction with INLA: application to IFU galaxy data

S González-Gaitán ✉, R S de Souza, A Krone-Martins, E Cameron, P Coelho,
L Galbany, E E O Ishida, COIN collaboration

Monthly Notices of the Royal Astronomical Society, Volume 482, Issue 3, 21
January 2019, Pages 3880–3891, <https://doi.org/10.1093/mnras/sty2881>

Spatially aware techniques
to better exploit spatial
information





Spatial field reconstruction with INLA: application to IFU galaxy data

S González-Gaitán ✉, R S de Souza, A Krone-Martins, E Cameron, P Coelho,
L Galbany, E E O Ishida, COIN collaboration

Monthly Notices of the Royal Astronomical Society, Volume 482, Issue 3, 21
January 2019, Pages 3880–3891, <https://doi.org/10.1093/mnras/sty2881>

The proposed Bayesian approach enables the recovery of areas with bad pixels, and an increased power to detect structures in data sets subject to substantial noise and/or sparsity of sampling.



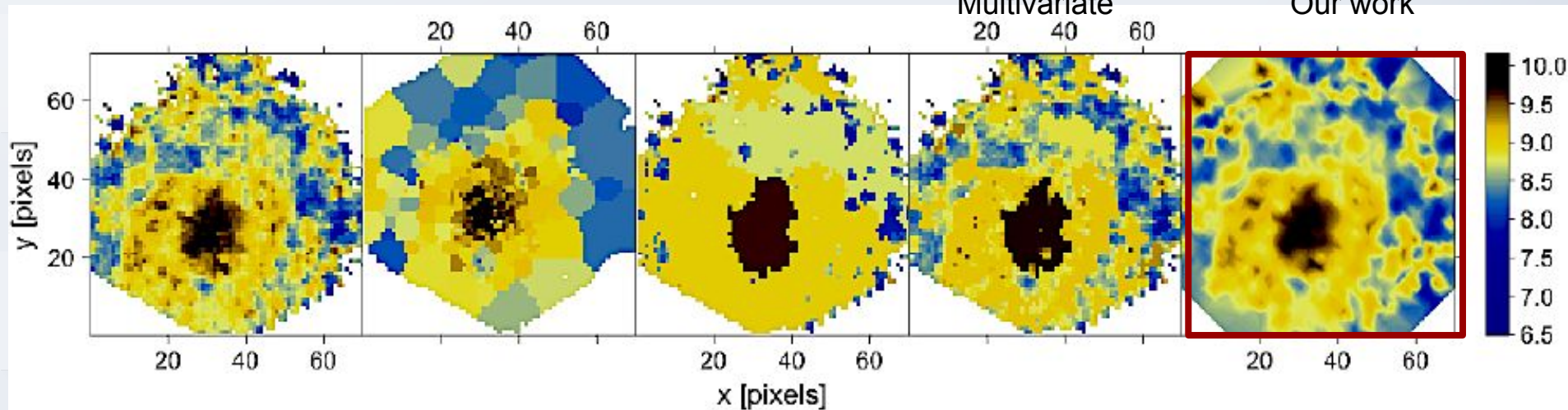
Input Data

Voronoi

Bayesian Voronoi

Multivariate

Our work





The Cosmostatistics Initiative

The Cosmostatistics Initiative (COIN) is an international network which aims to create an interdisciplinary environment where collaborations between astronomers, statisticians and machine learning experts can flourish. The group utilizes a management model which can find parallel in technological start-ups: based on a dynamic, non-hierarchical and people-centric approach.

The LSST Dark Energy Science Collaboration

The LSST Dark Energy Science Collaboration (DESC) is an international collaboration preparing for a variety of cosmological analyses with the Large Synoptic Survey Telescope (LSST) data. In advance of LSST's first observations, DESC will help prepare for LSST science analysis, make synergistic connections with ongoing cosmological surveys and provide the dark energy community with state of the art analysis tools.



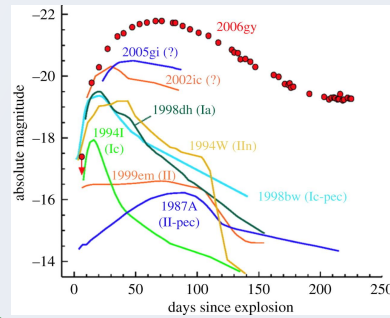
RESSPECT

The REcommendation System for SPECTroscopic follow-up (RESSPECT) is a collaboration between COIN and LSST-DESC which aims to adapt active learning strategies for the construction of optimized training samples for supernova photometric classification in the context of LSST.

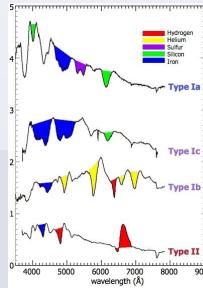
The team is formed by researchers from both collaborations who are working together in the development of a recommendation system which will enable informed decisions regarding the allocation of spectroscopic follow-up resources and consequent optimized scientific results from purely photometric samples.

Spectroscopy (high-resolution) vs Photometry (low-resolution)

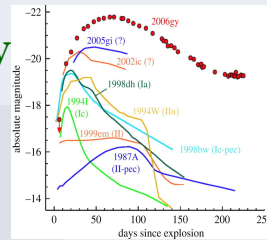
Apply it here



Learn from here



Photometry
+
Spectra

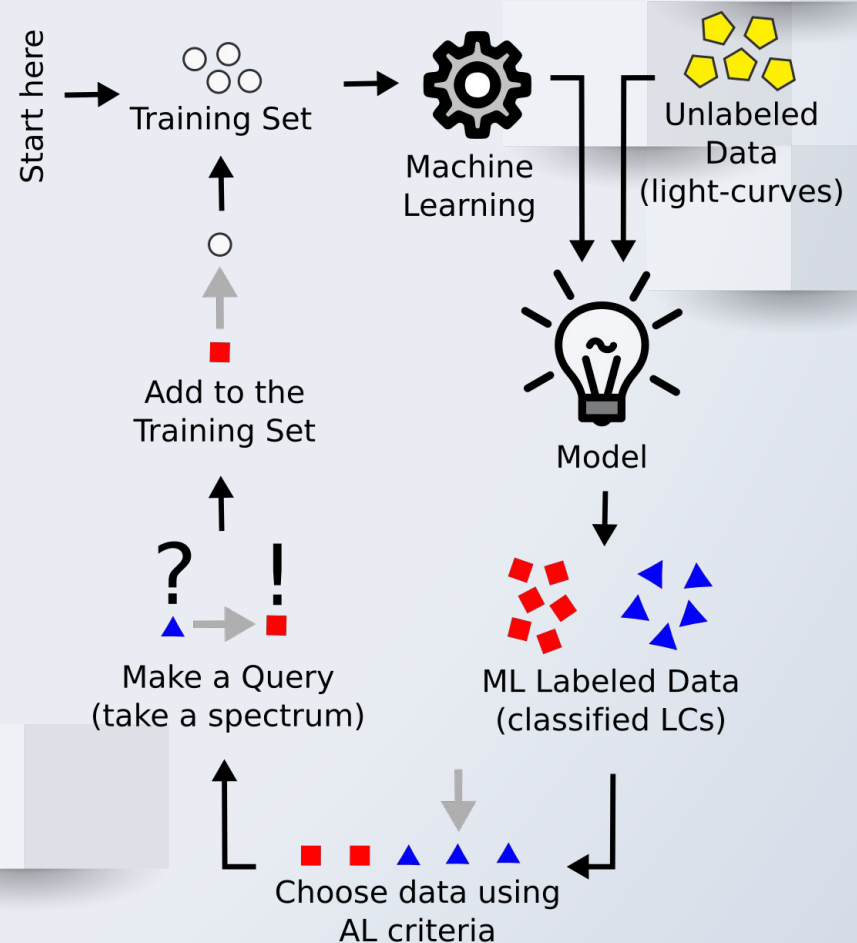
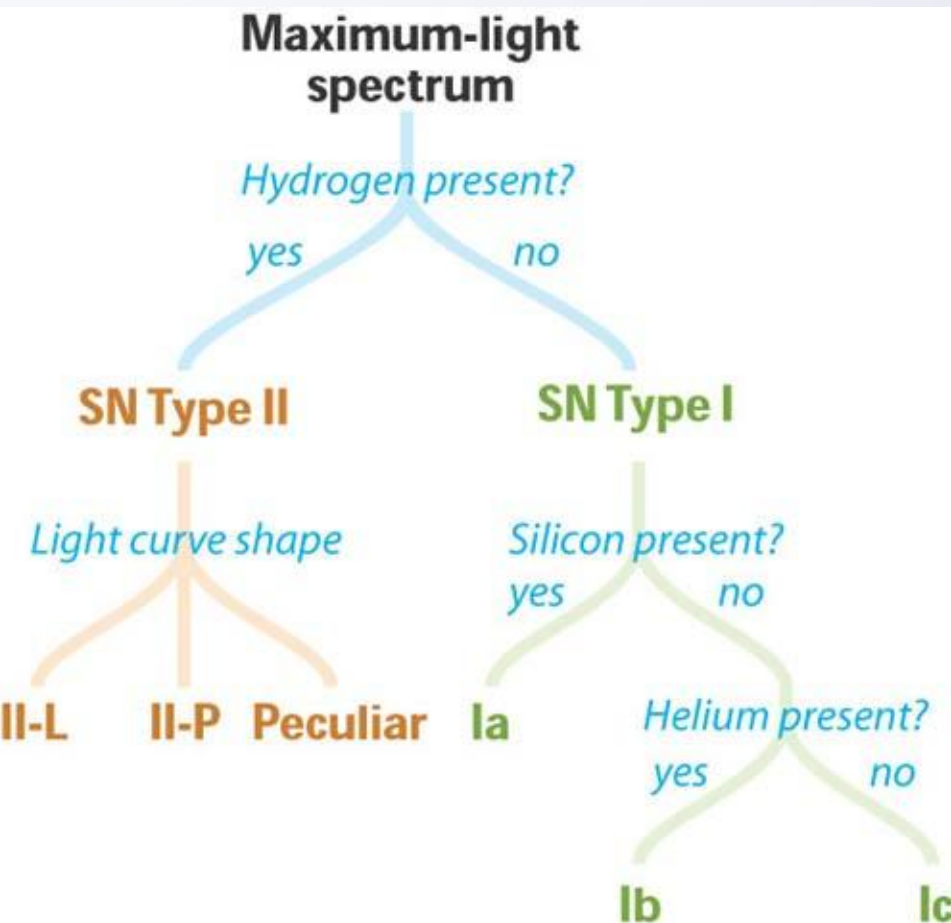


Photometry
only

Optimizing spectroscopic follow-up strategies for supernova photometric classification with active learning

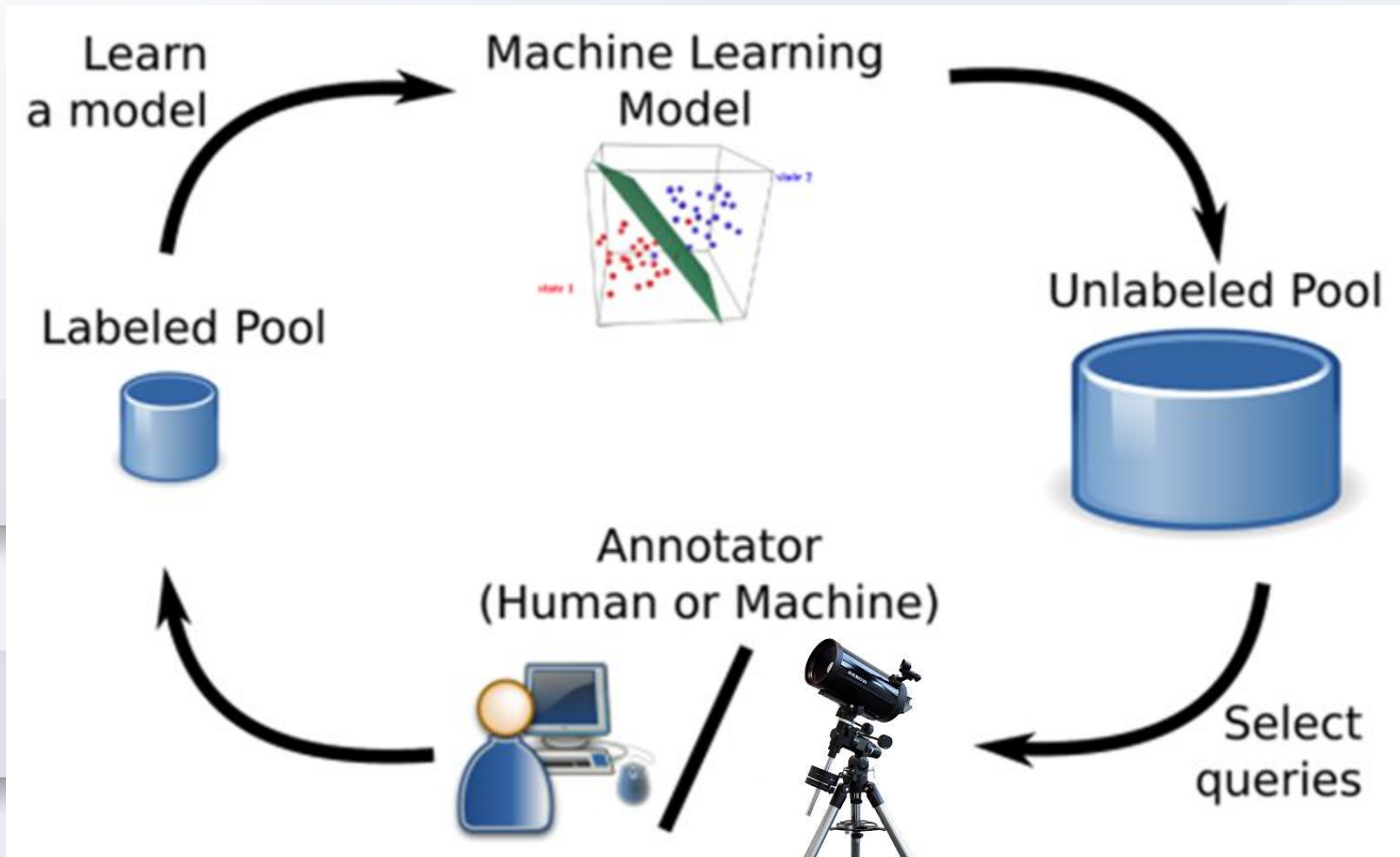
E E O Ishida ✉, R Beck, S González-Gaitán, R S de Souza, A Krone-Martins, J W Barrett, N Kennamer, R Vilalta, J M Burgess, B Quint, ... [Show more](#)

Monthly Notices of the Royal Astronomical Society, Volume 483, Issue 1, 11 February 2019, Pages 2–18,



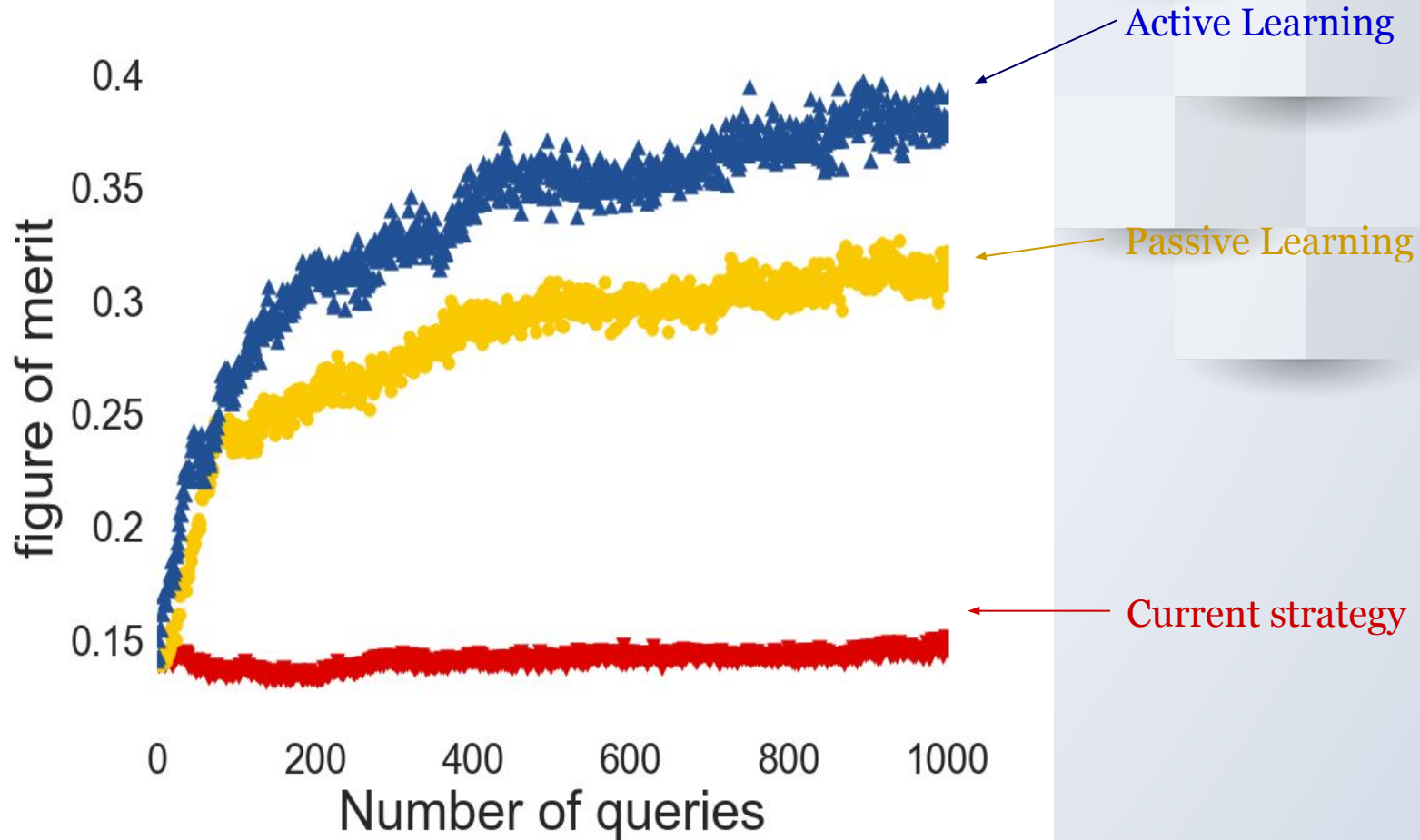
Active Learning

Optimal classification, minimum training



AL for Supernova Classification

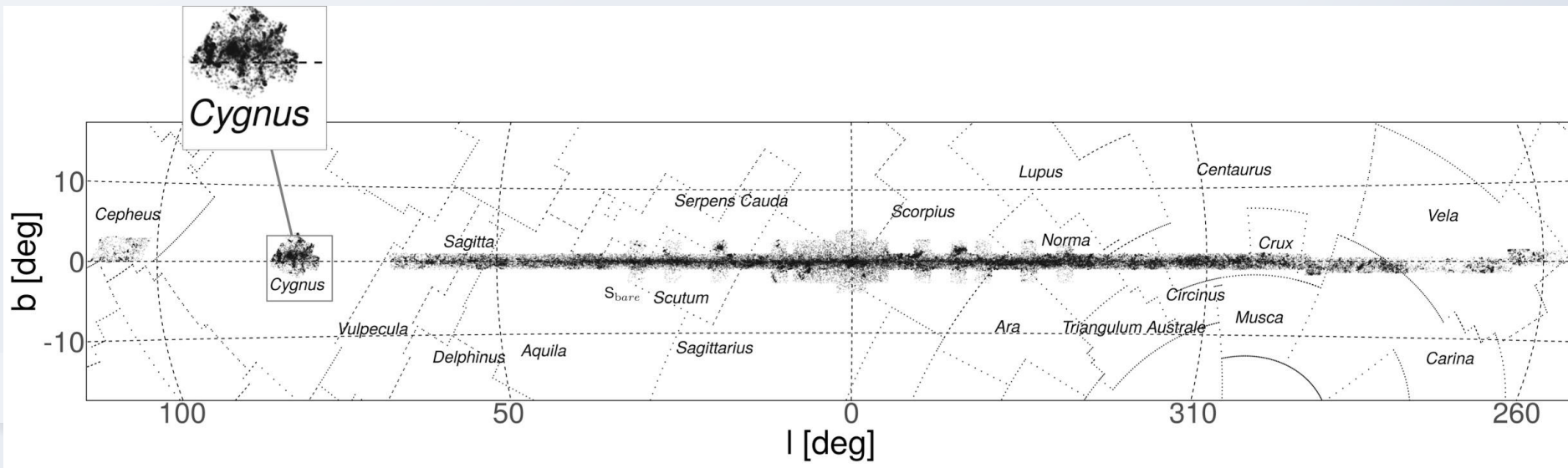
Diagnostic plots

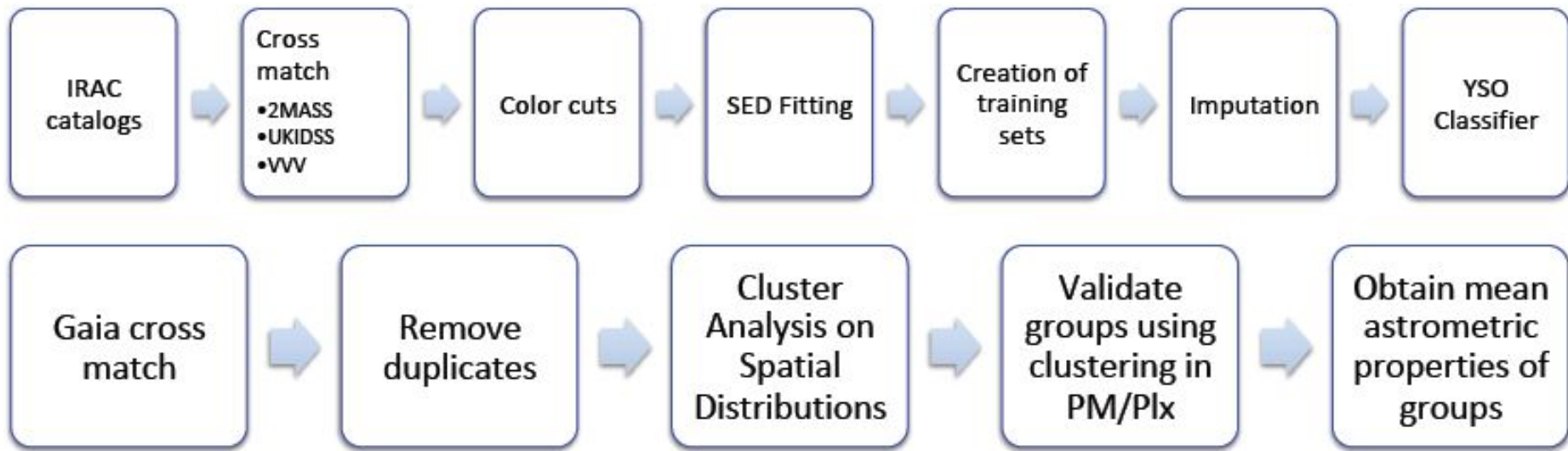


CYMES - Candidate YSOs with Mid-IR Excess from Spitzer

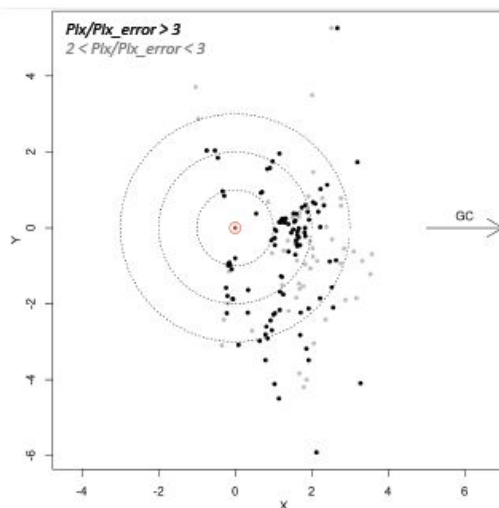
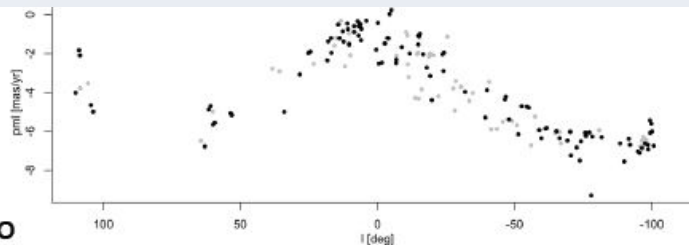
We perform an automated search for young stellar objects (YSOs) using the public IRAC catalogs from Spitzer legacy surveys (GLIMPSE I/II, 3D, Vela-Carina, Cygnus X, and SMOG). They are augmented by near-infrared (NIR) catalogs from 2MASS, UKIDSS, and VVV.

We identify ~150,000 YSOs candidates.
More than twice the previously known objects.





- About one third of YSO candidates have *Gaia* astrometry
- We expect YSOs that are members of associations to cluster in parallax/proper-motions space, while contaminants are not expected to be clustered
- *Gaia* can be used to validate which groups are real and which YSO candidates are likely members of these groups



Mean proper motions (top) and distances (left) for groups of stars clustered in pos-plx-pm space.

Estimating Contamination Rate for YSO Candidates in the Galactic Plane

Michael A. Kuhn (CIT), Lynne Hillenbrand (CIT), Alberto Krone-Martins (UC Irvine),
Rafael S. de Souza (SHAO), Emille E. O. Ishida (CNRS), Alfred Castro-Ginard (U Barcelona)

We request three nights for observations with the P200/DBSP to spectroscopically classify a sample of candidate young stellar objects (YSOs).

Some potential collaboration with SHAO to use local facilities?

2020B Palomar Cover Sheet

4/15/20, 4:41 PM



COO/Palomar Observatory
Cover Sheet - Semester 2020B
200-inch Telescope

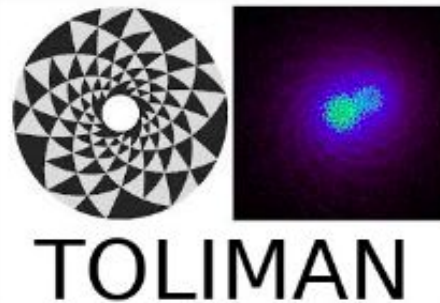




COIN-Focus

Toliman

The TOLIMAN space telescope is a low-cost, agile mission concept dedicated to astrometric detection of exoplanets in the near-solar environment, and particularly targeting the Alpha Cen system.



CRP#7?



- ❑ COIN was constructed under the lemma *people come first*.
- ❑ The scientific project emerges from a shared group interest.
- ❑ They are a product of the interaction between a unique group of people, whose materialisation is only possible in an environment which profoundly respects the diversity of their scientific backgrounds, gender, career stages and nationalities.



THANK
YOU

